Gravity Equations 03

$$a = \frac{v^2}{R} \qquad \qquad v = \frac{2\pi R}{T}$$

$$F = m \bullet a$$
 $F = m \bullet g$

$$F_g = \frac{G \bullet M \bullet m}{R^2}$$

$$\frac{G \cdot M \cdot m}{R^2} = \frac{4 \cdot m \cdot \pi^2 \cdot R}{T^2}$$

$$\frac{R^3}{T^2} = K \qquad Kepler's \ Law$$

 $v = \sqrt{\frac{G \cdot M}{R}}$ satellite speed (circular orbit)

$$E_p = -\frac{G \cdot M \cdot m}{R}$$
 gravitational potential energy

 $W = \Delta E_p$ work = change in potential energy

$$v_{escape} = \sqrt{\frac{2 \cdot G \cdot M}{R}}$$
 escape speed

- $E_t = E_k + E_p$ total energy = kinetic energy + potential energy
- $E_{t1} = E_{t2}$ total energy is conserved